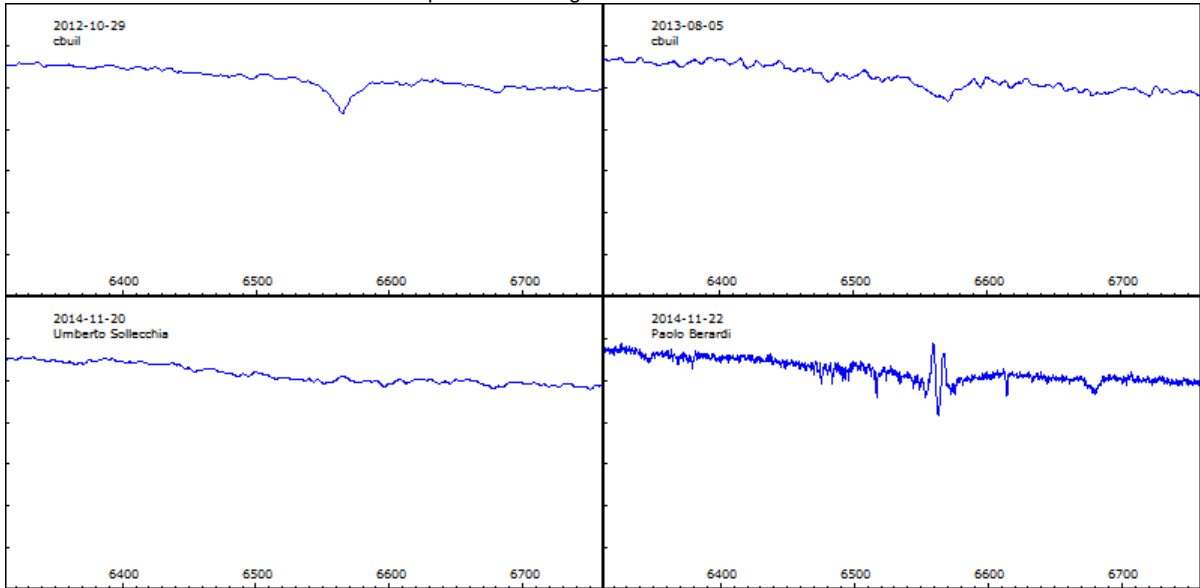


Emission increase since last observations

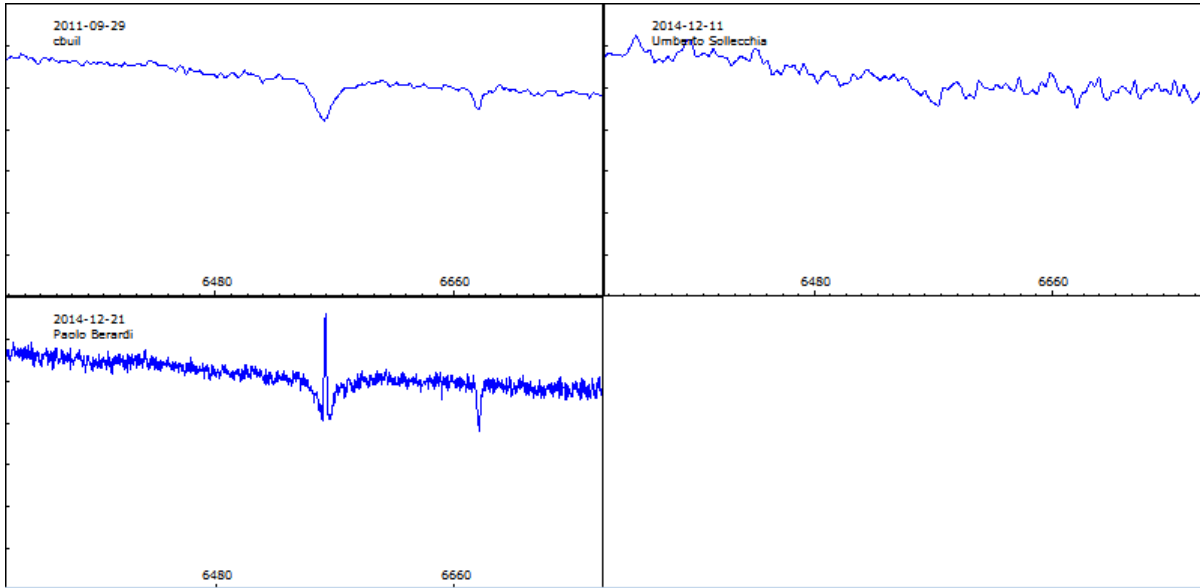
BD+62 11

Emission phase restarting – detected at resolution R 500



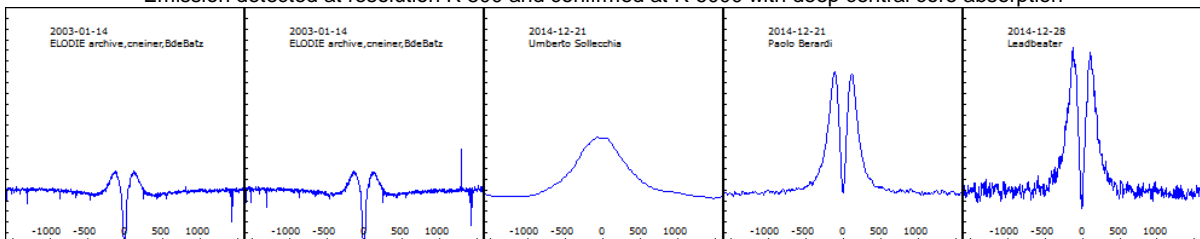
V548 per

Emission detected at resolution R 500



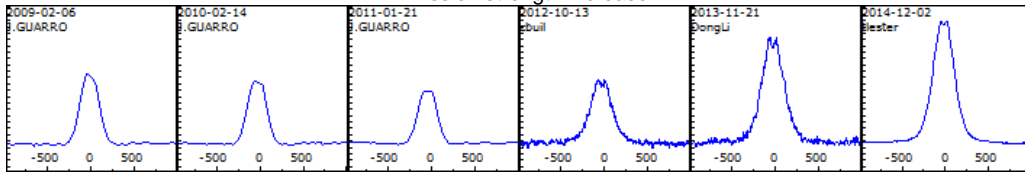
BD+55 521

Emission detected at resolution R 500 and confirmed at R 6000 with deep central core absorption



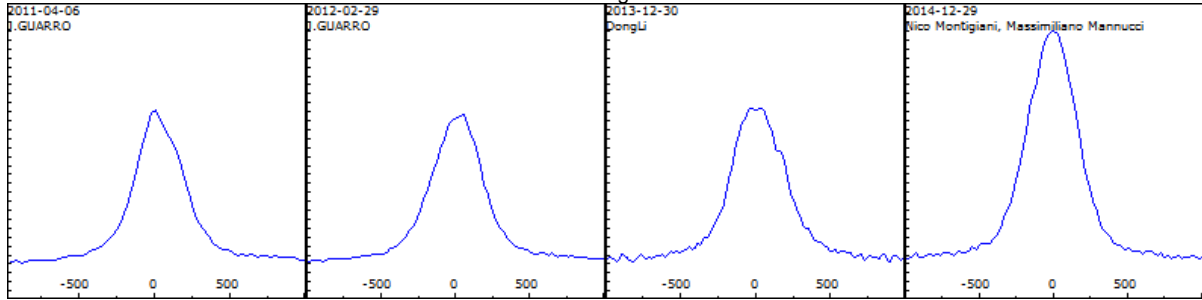
HD 23982

Emission strength increase



HD 245310

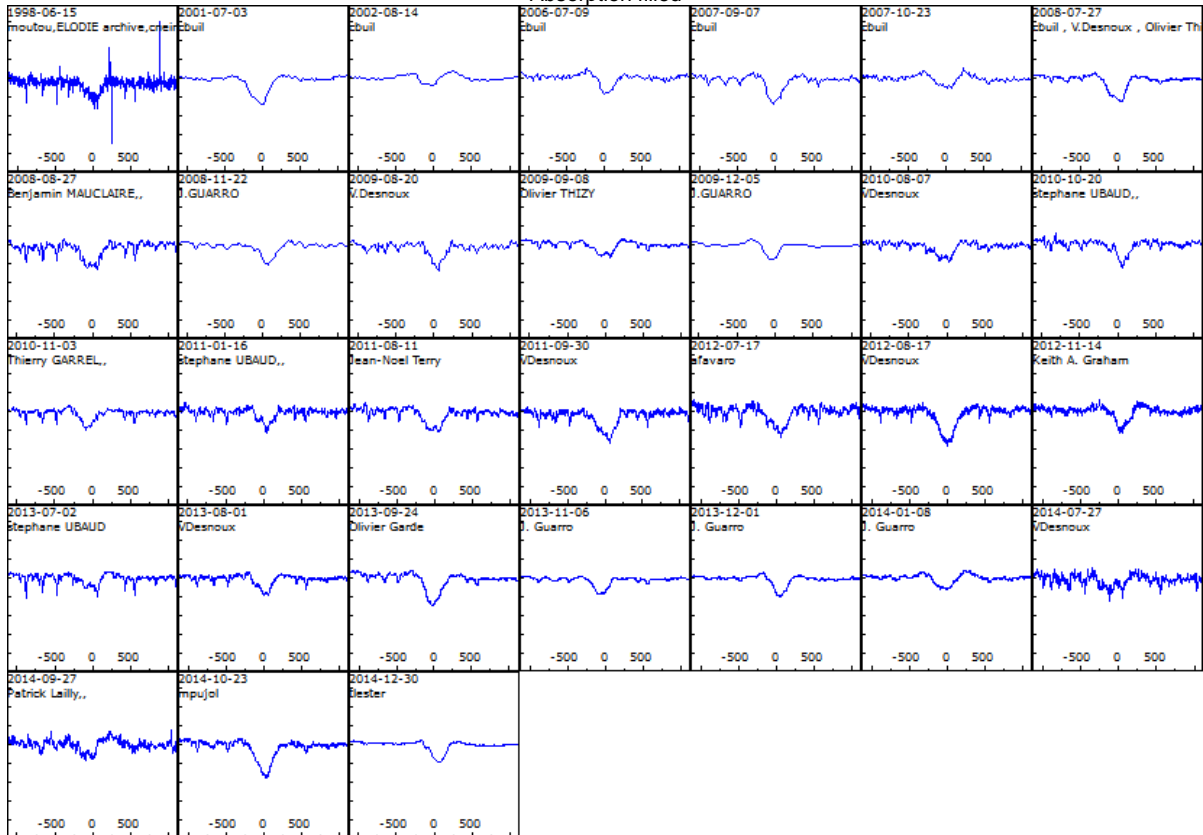
Emission strength increase



Moderate evolutions of H-alpha line

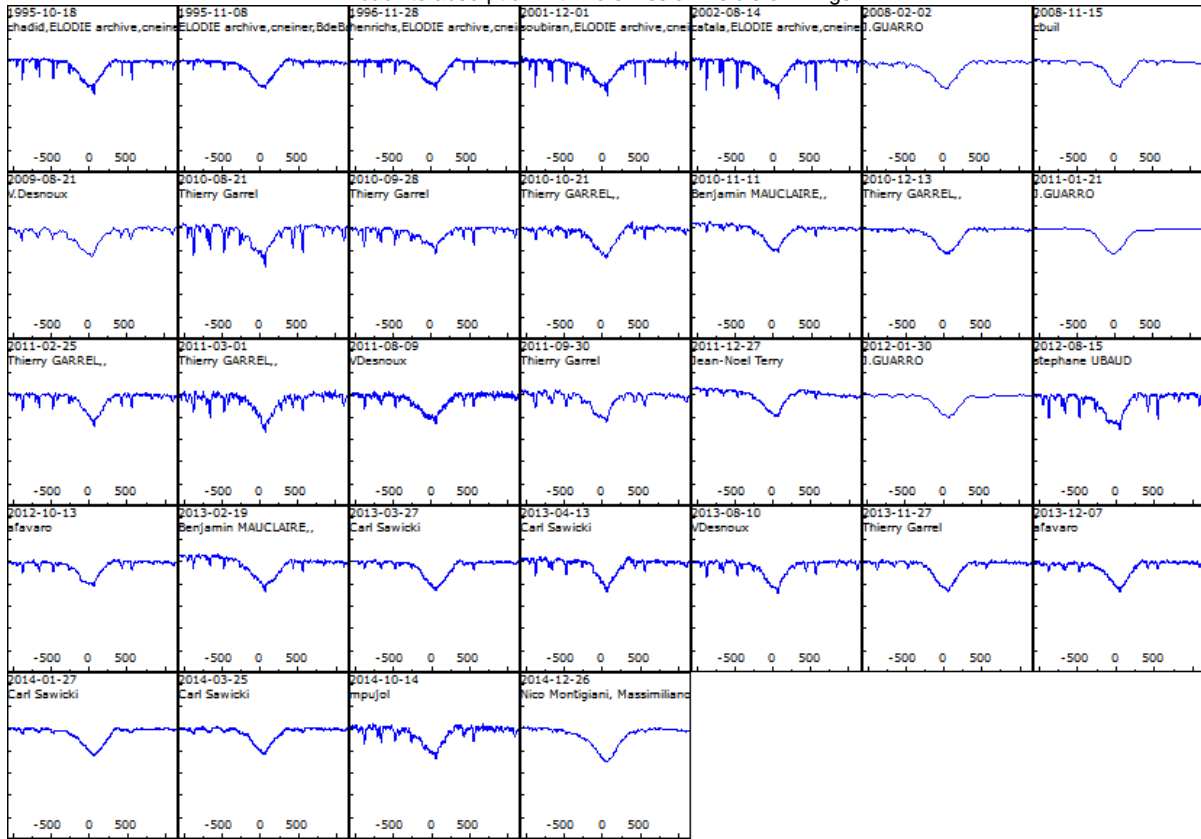
14 lac

Absorption filled



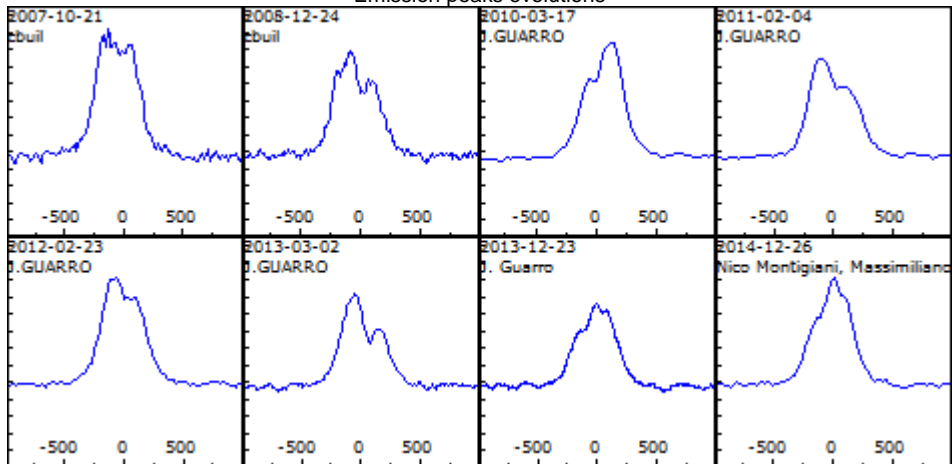
ksi Per

back to absorption with no emission visible on wings



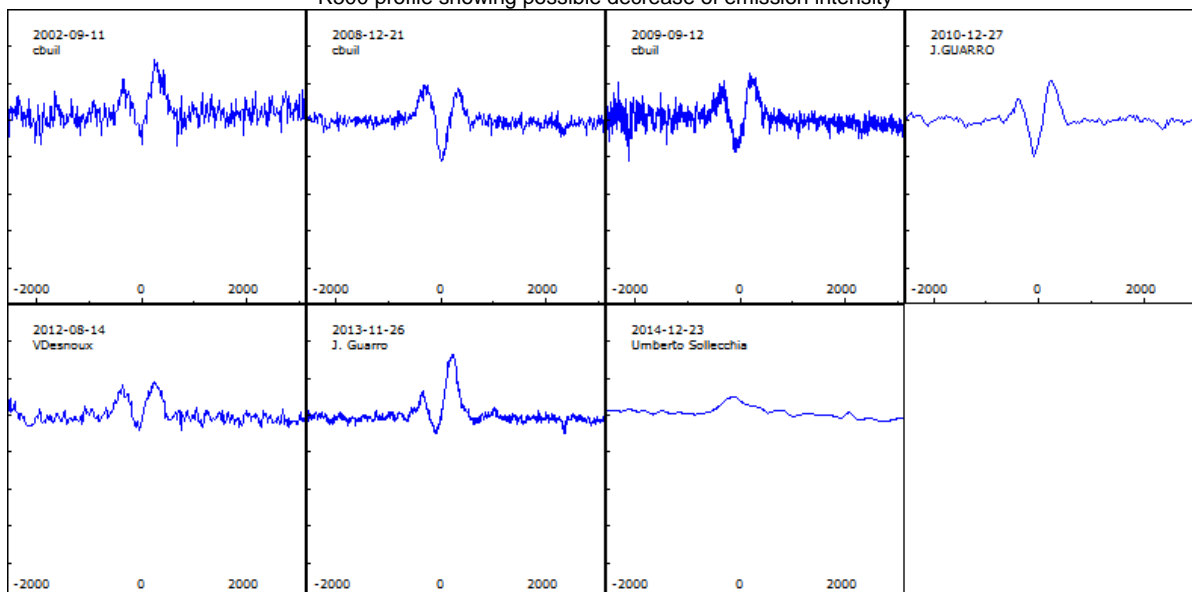
V415 Aur

Emission peaks evolutions



V473 Per

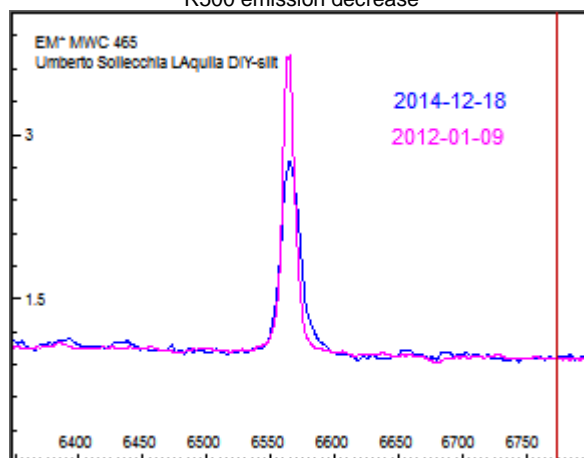
R500 profile showing possible decrease of emission intensity



Emission decrease of H-alpha line

EM* MWC 563

R500 emission decrease

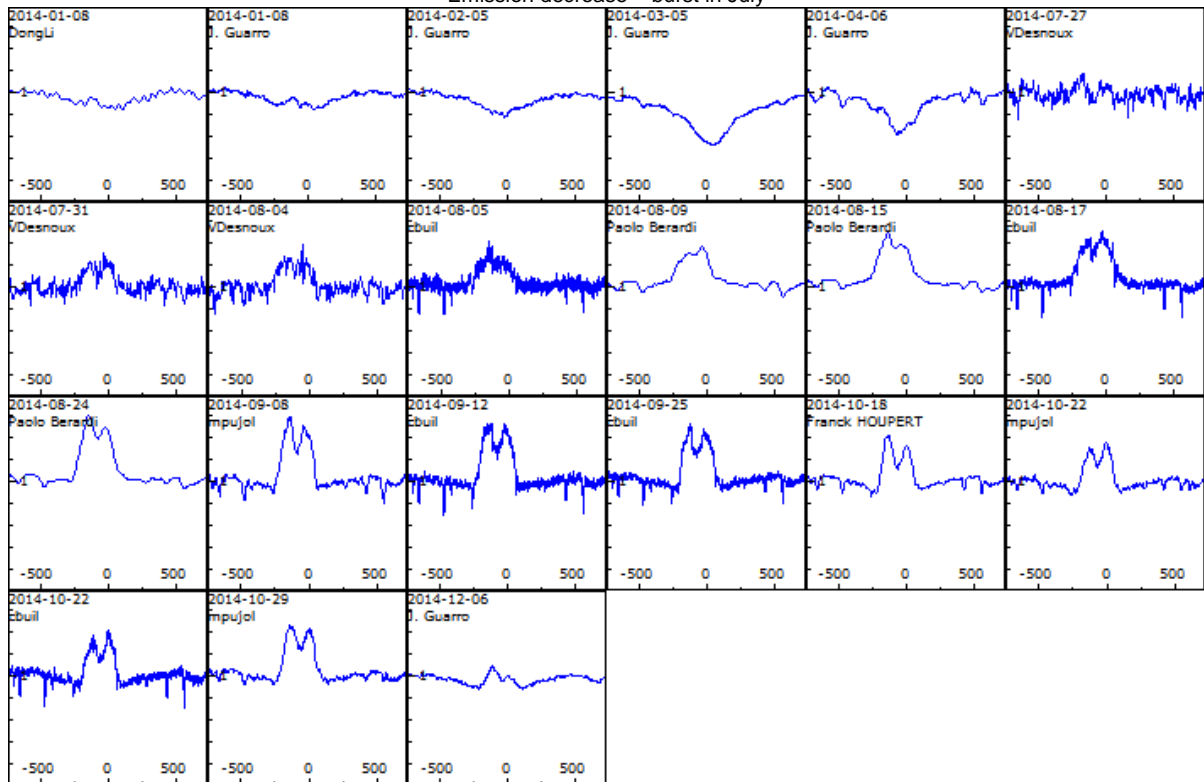


V438 Aur

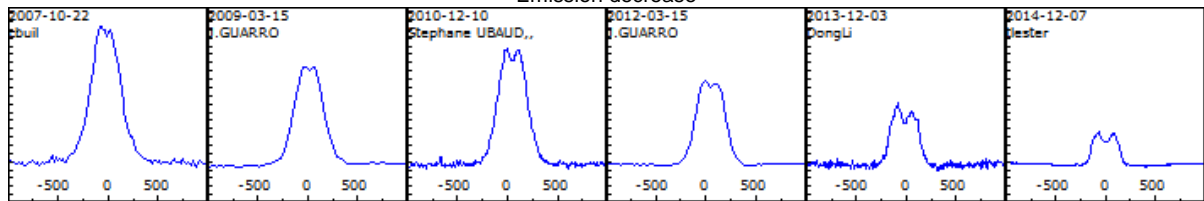
Emission wings below continuum



V442 And
Emission decrease – burst in July



V447 Aur
Emission decrease

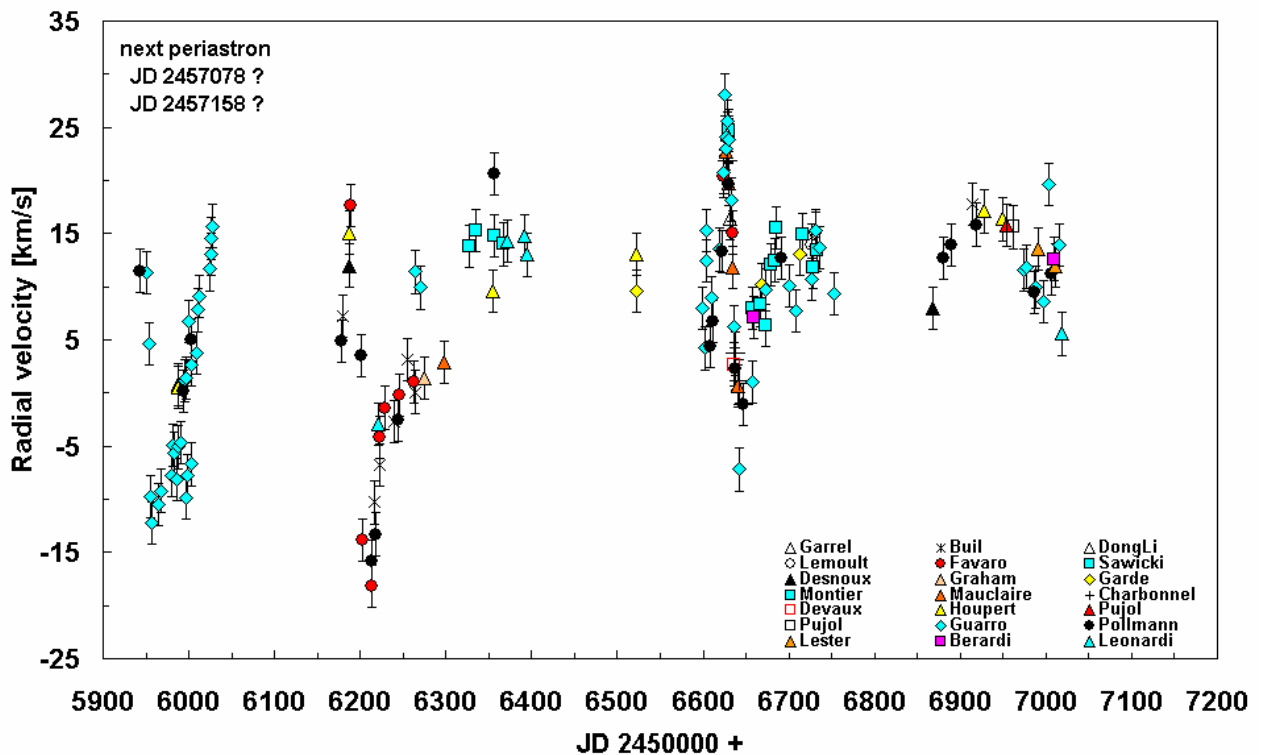
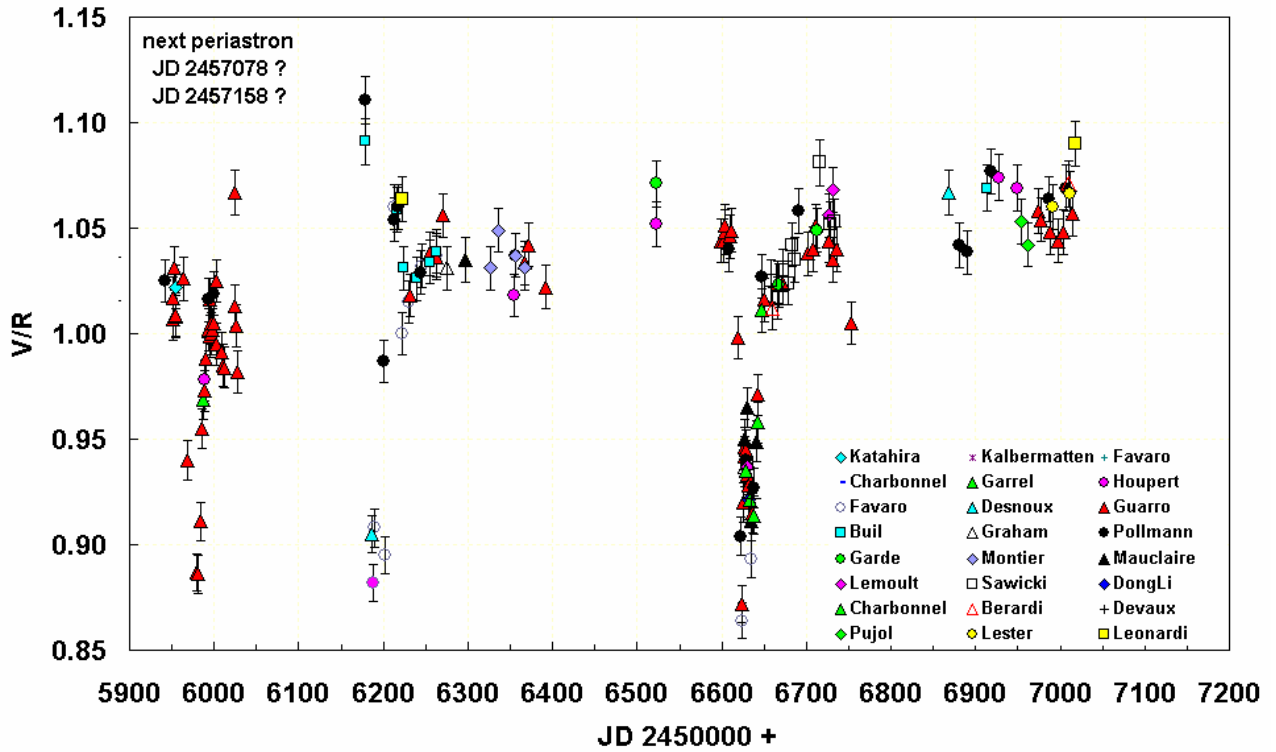


Be monitoring projects

By Ernst Pollmann

28 Tau

Based on our V/R and RV monitoring of the Be binary 28 Tau, its known orbital period of 218 days, the observed periastron at JD 2455955, 2456213 & 2456642 (the following Fig.) we should expect the next periastron at approx. JD 2457078. The second page shows the updated relationship between H α EW and V magnitude in different phases (Be & Be shell).



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